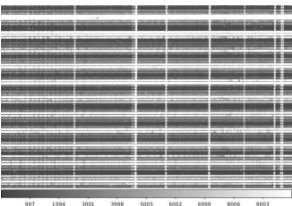


# machine learning for stellar spectra

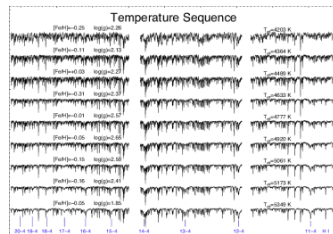
Sébastien Fabbro

*NTCO 2017*

raw 2D image



calibrated 1D spectra



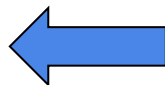
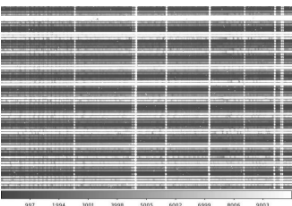
estimated stellar parameters

$T_{\text{eff}}$ ,  $\log(g)$ ,  $[\text{Fe}/\text{H}]$ ,  
 $[\text{C}/\text{H}]$ ,  $[\text{N}/\text{H}]$ ,  $[\text{O}/\text{H}]$ ,  
 $[\text{Na}/\text{H}]$ ,  $[\text{Mg}/\text{H}]$ ,  $[\text{Al}/\text{H}]$ ,  
 $[\text{Si}/\text{H}]$ ,  $[\text{S}/\text{H}]$ ,  $[\text{K}/\text{H}]$ ,  
 $[\text{Ca}/\text{H}]$ ,  $[\text{Ti}/\text{H}]$ ,  $[\text{V}/\text{H}]$ ,  
 $[\text{Mn}/\text{H}]$ ,  $[\text{Ni}/\text{H}]$ ,...

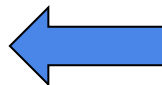
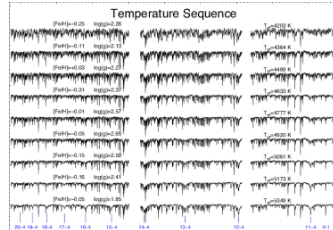


stellar populations  
galactic archeology  
near-field cosmology

synthetic 2D image

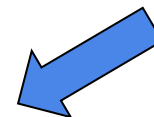


synthetic 1D spectra

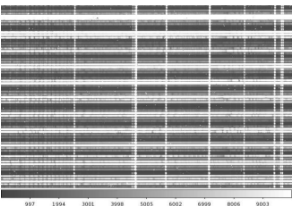


stellar parameters

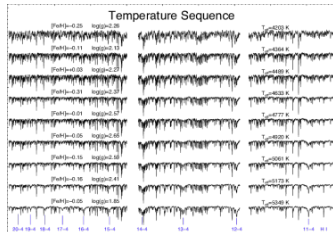
$T_{\text{eff}}$ ,  $\log(g)$ ,  $[\text{Fe}/\text{H}]$ ,  
 $[\text{C}/\text{H}]$ ,  $[\text{N}/\text{H}]$ ,  $[\text{O}/\text{H}]$ ,  
 $[\text{Na}/\text{H}]$ ,  $[\text{Mg}/\text{H}]$ ,  $[\text{Al}/\text{H}]$ ,  
 $[\text{Si}/\text{H}]$ ,  $[\text{S}/\text{H}]$ ,  $[\text{K}/\text{H}]$ ,  
 $[\text{Ca}/\text{H}]$ ,  $[\text{Ti}/\text{H}]$ ,  $[\text{V}/\text{H}]$ ,  
 $[\text{Mn}/\text{H}]$ ,  $[\text{Ni}/\text{H}]$ ,...



raw 2D image



calibrated 1D spectra



estimated stellar parameters

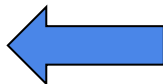
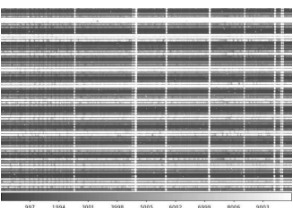
$T_{\text{eff}}$ ,  $\log(g)$ ,  $[\text{Fe}/\text{H}]$ ,  
 $[\text{C}/\text{H}]$ ,  $[\text{N}/\text{H}]$ ,  $[\text{O}/\text{H}]$ ,  
 $[\text{Na}/\text{H}]$ ,  $[\text{Mg}/\text{H}]$ ,  $[\text{Al}/\text{H}]$ ,  
 $[\text{Si}/\text{H}]$ ,  $[\text{S}/\text{H}]$ ,  $[\text{K}/\text{H}]$ ,  
 $[\text{Ca}/\text{H}]$ ,  $[\text{Ti}/\text{H}]$ ,  $[\text{V}/\text{H}]$ ,  
 $[\text{Mn}/\text{H}]$ ,  $[\text{Ni}/\text{H}]$ ,...



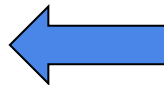
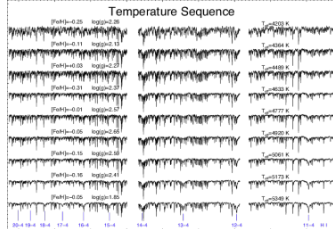
typical ML  
analysis

stellar populations  
galactic archeology  
near-field cosmology

synthetic 2D image

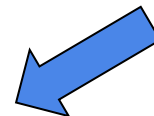


synthetic 1D spectra



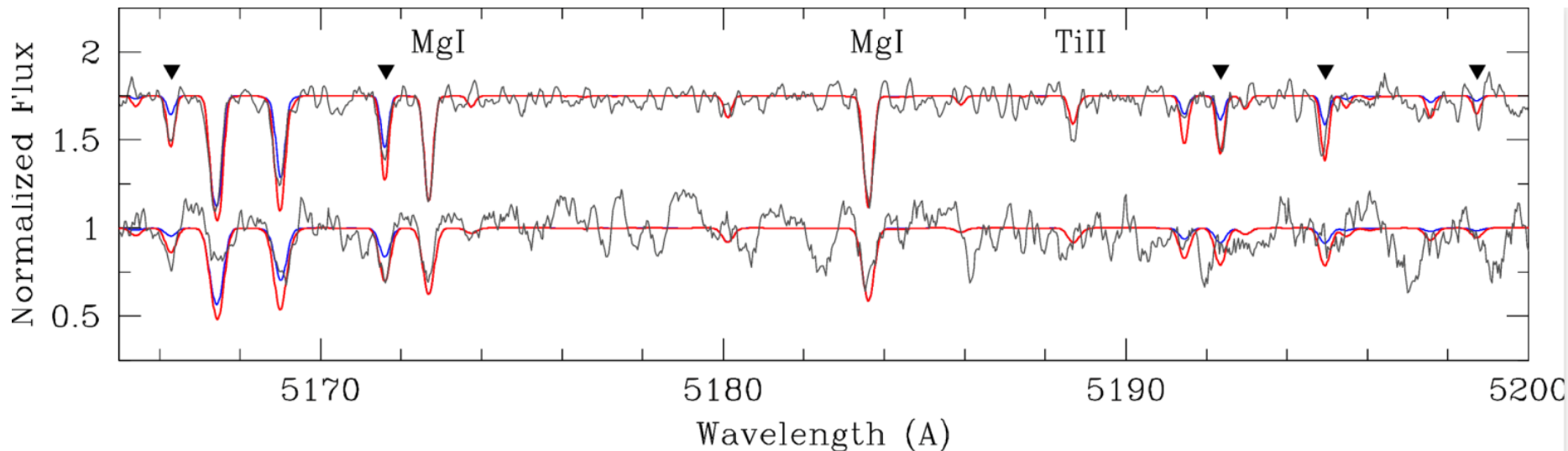
stellar parameters

$T_{\text{eff}}$ ,  $\log(g)$ ,  $[\text{Fe}/\text{H}]$ ,  
 $[\text{C}/\text{H}]$ ,  $[\text{N}/\text{H}]$ ,  $[\text{O}/\text{H}]$ ,  
 $[\text{Na}/\text{H}]$ ,  $[\text{Mg}/\text{H}]$ ,  $[\text{Al}/\text{H}]$ ,  
 $[\text{Si}/\text{H}]$ ,  $[\text{S}/\text{H}]$ ,  $[\text{K}/\text{H}]$ ,  
 $[\text{Ca}/\text{H}]$ ,  $[\text{Ti}/\text{H}]$ ,  $[\text{V}/\text{H}]$ ,  
 $[\text{Mn}/\text{H}]$ ,  $[\text{Ni}/\text{H}]$ ,...



# traditional approach

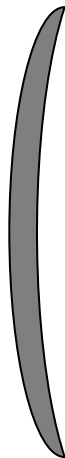
Gemini-GRACES spectroscopy, Venn et. al (2017)

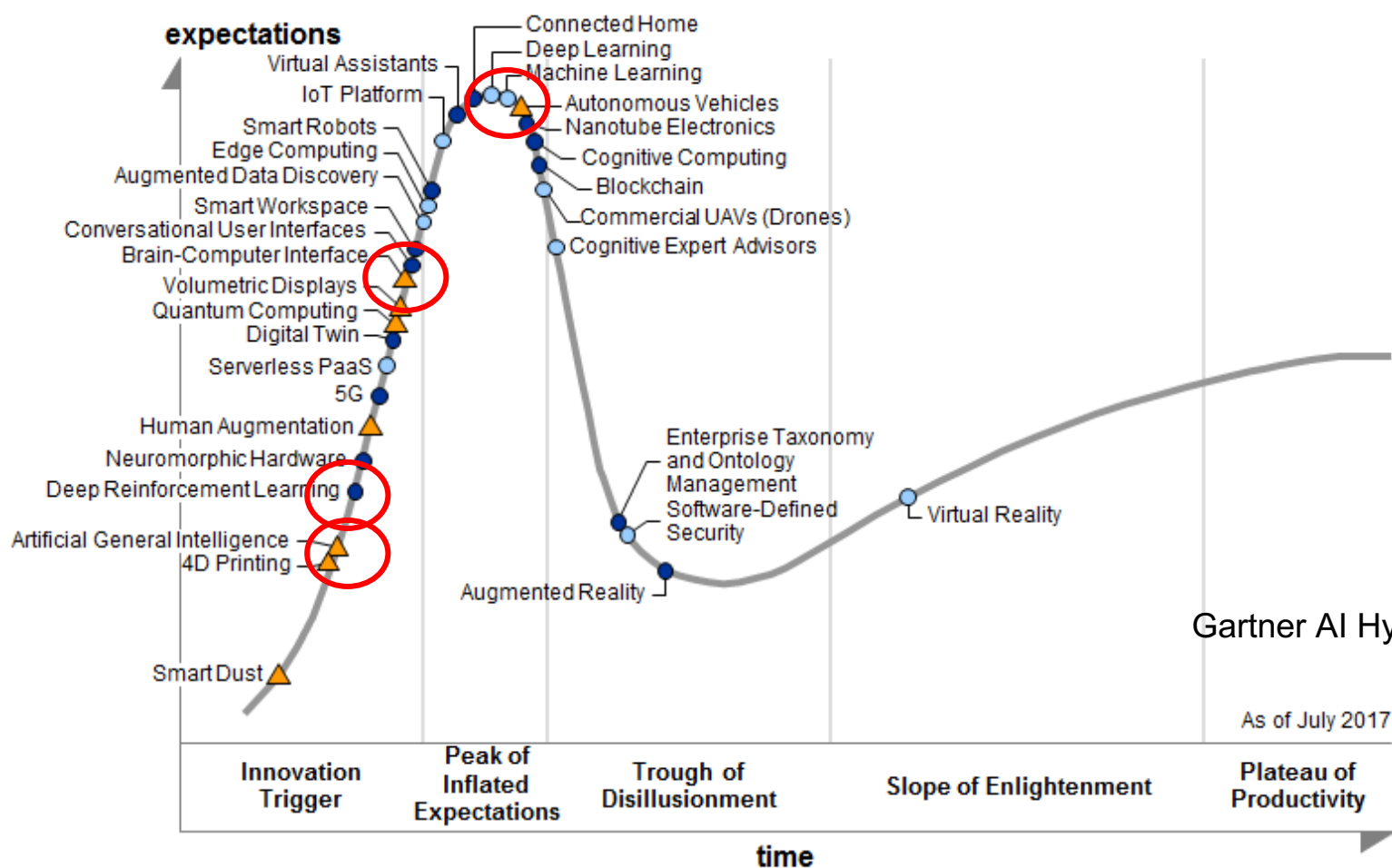


- physics models methods (EQW, NL fit, projections...)
- "data driven" methods (The Cannon, ANN): need a set of reference stars



AI





Gartner AI Hype Cycle

As of July 2017

**Years to mainstream adoption:**

○ less than 2 years

● 2 to 5 years

● 5 to 10 years

▲ more than 10 years

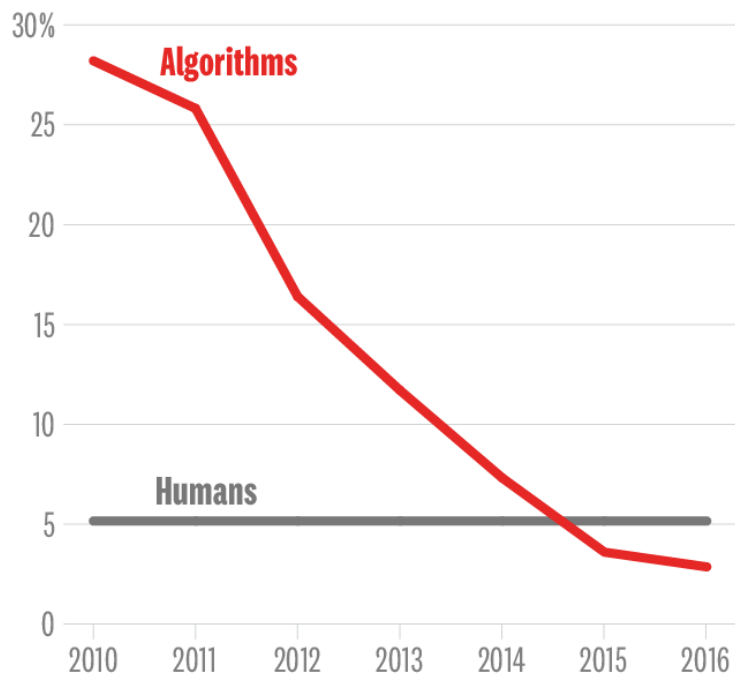
obsolete

⊗ before plateau

# recognize puppy from muffins



VISION ERROR RATE



SOURCE ELECTRONIC FRONTIER FOUNDATION

© HBR.ORG

**translate from 100 languages**

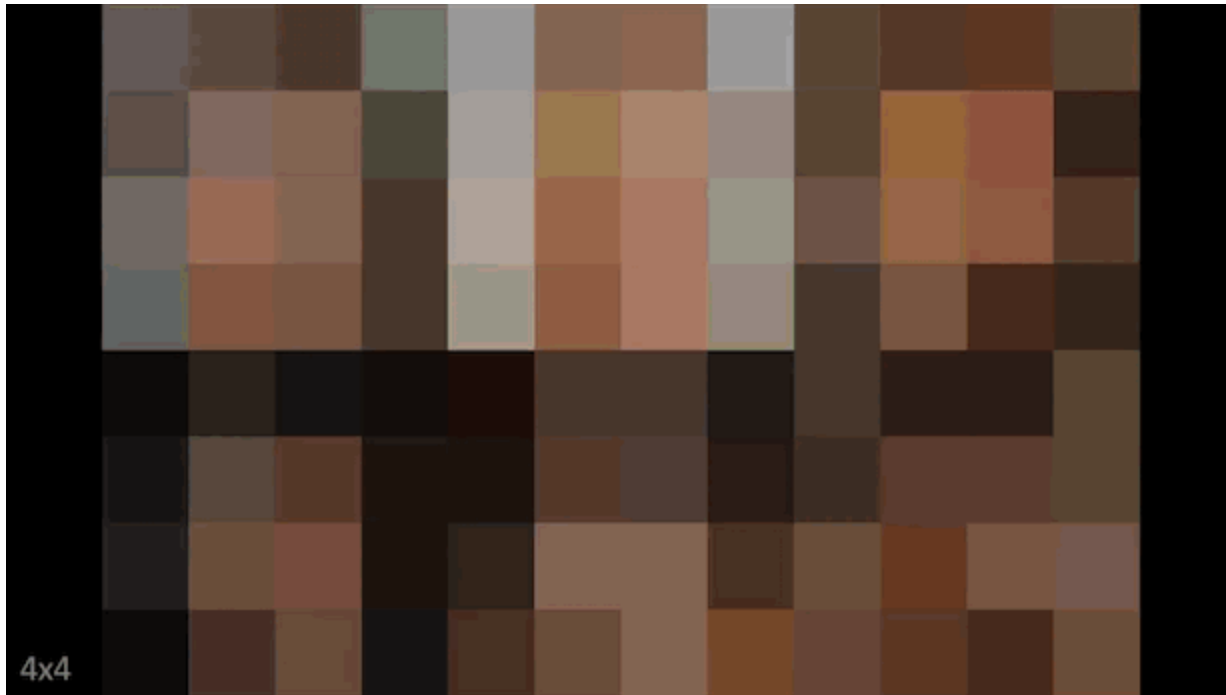




# restyle paintings



# generate new celebrities



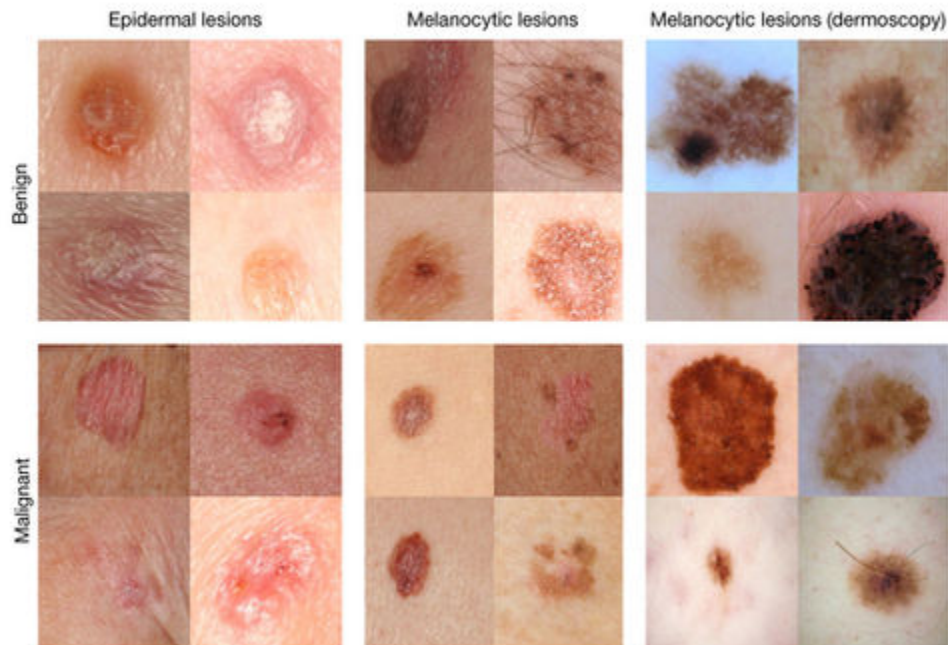
Karas et. al (2017)

# detect skin cancer

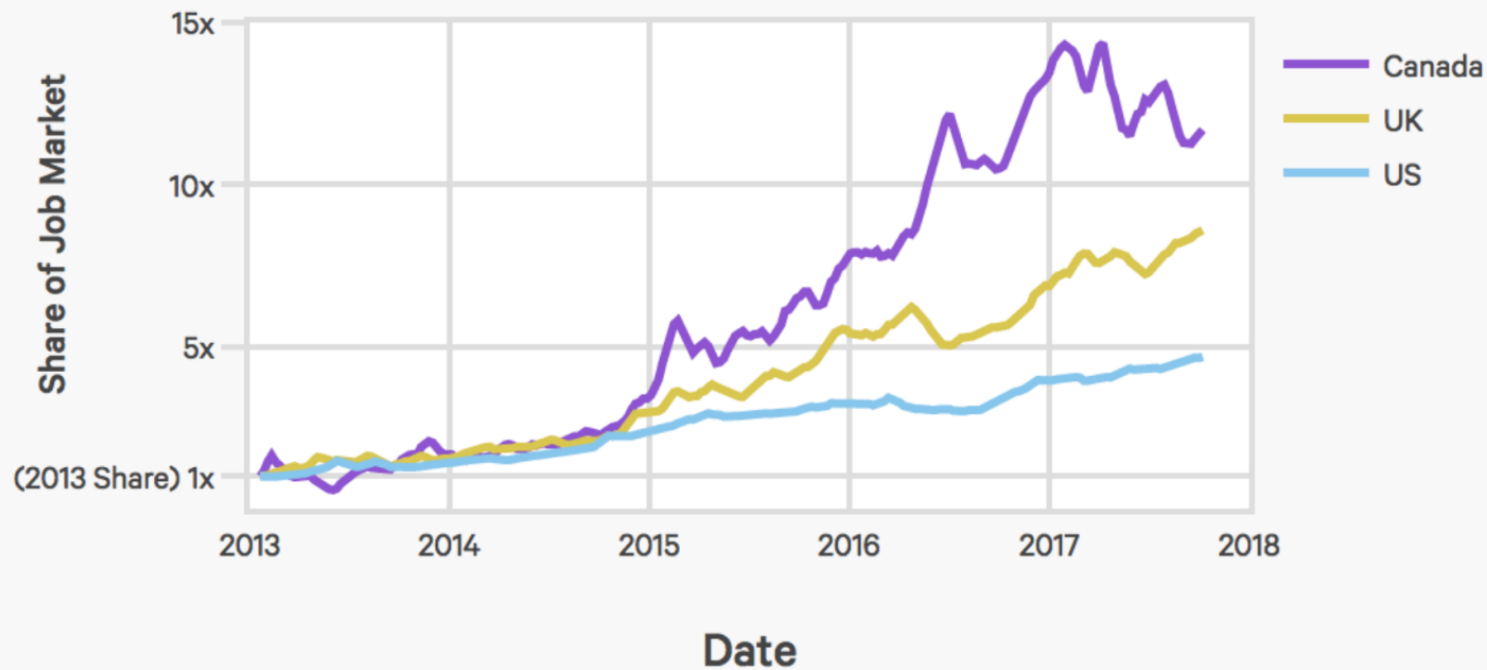
**a**



**b**

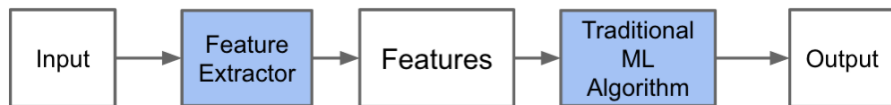


## Share of Jobs Requiring AI Skills (Indeed.com)

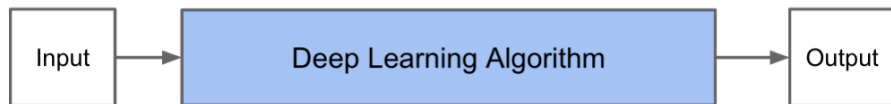




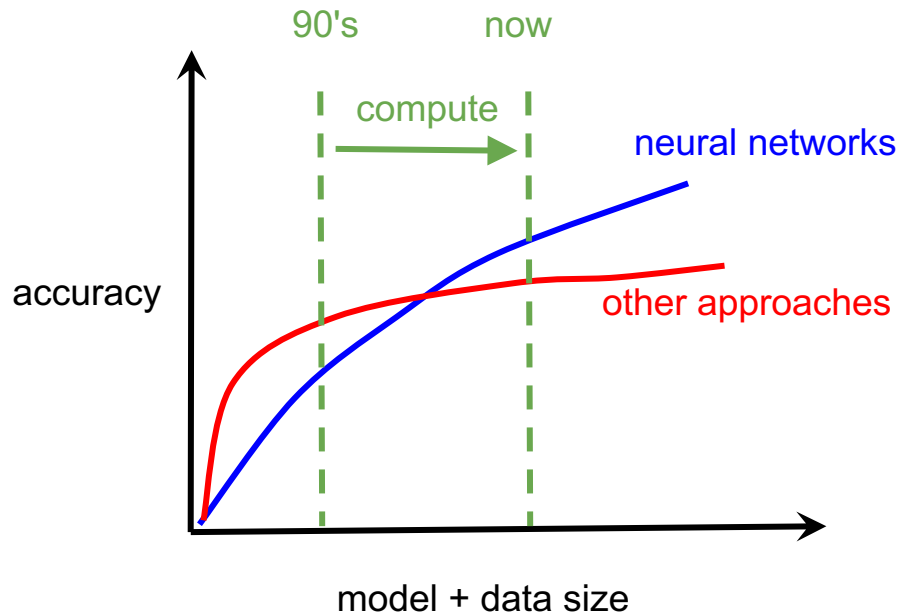
# deep learning



Traditional Machine Learning Flow



Deep Learning Flow

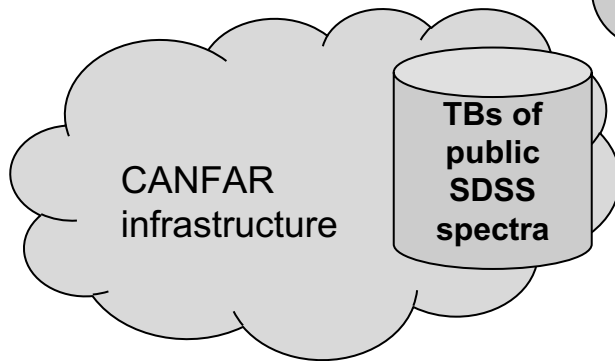


# should astronomers care?

- detect complex structures in large data sets
- classify astronomical objects
- learn time consuming simulations
- automate and accelerate manual data analysis tasks
- replace many image processing techniques
- powerful well written software
- prepare students

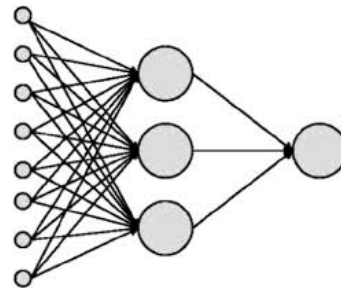


# an experiment



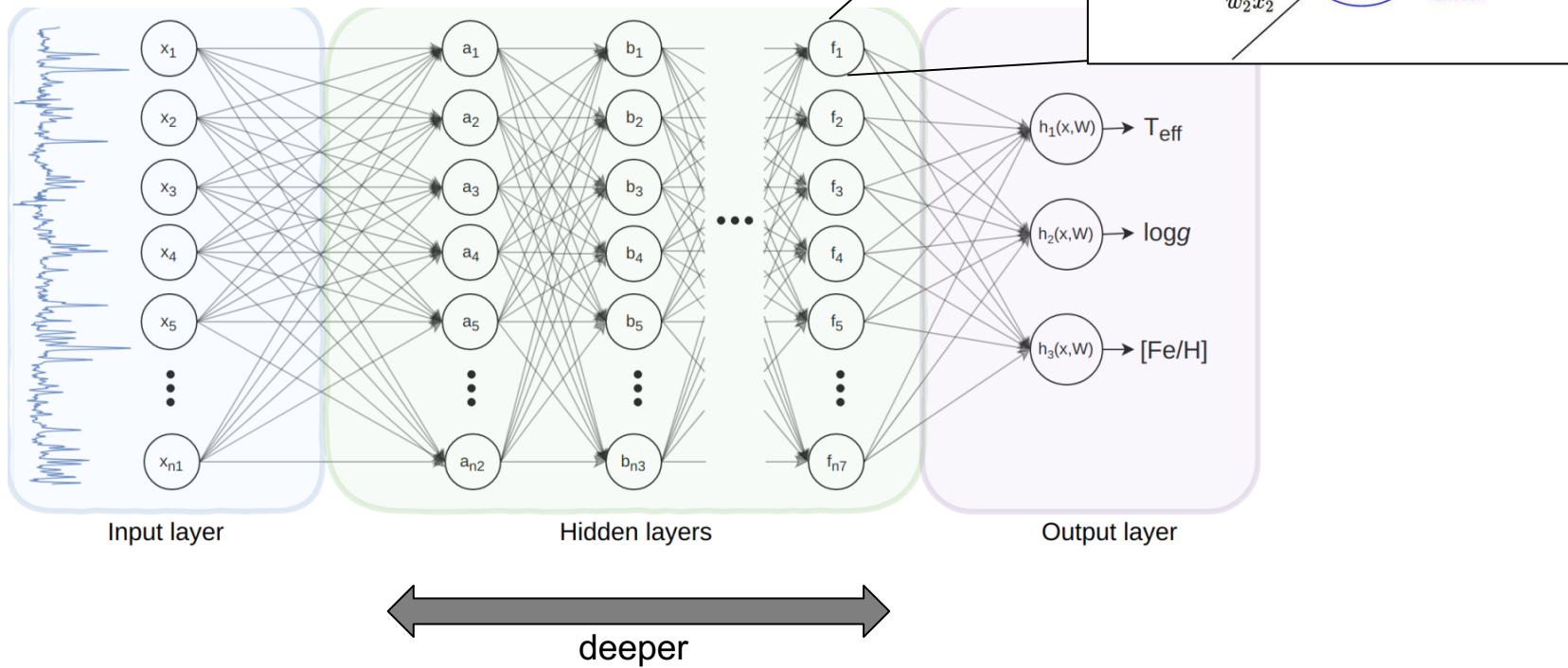
google: tensorflow, keras  
facebook: caffe2, pytorch  
amazon: mxnet, gluon  
microsoft: CNTK

**many choice of open DL frameworks**

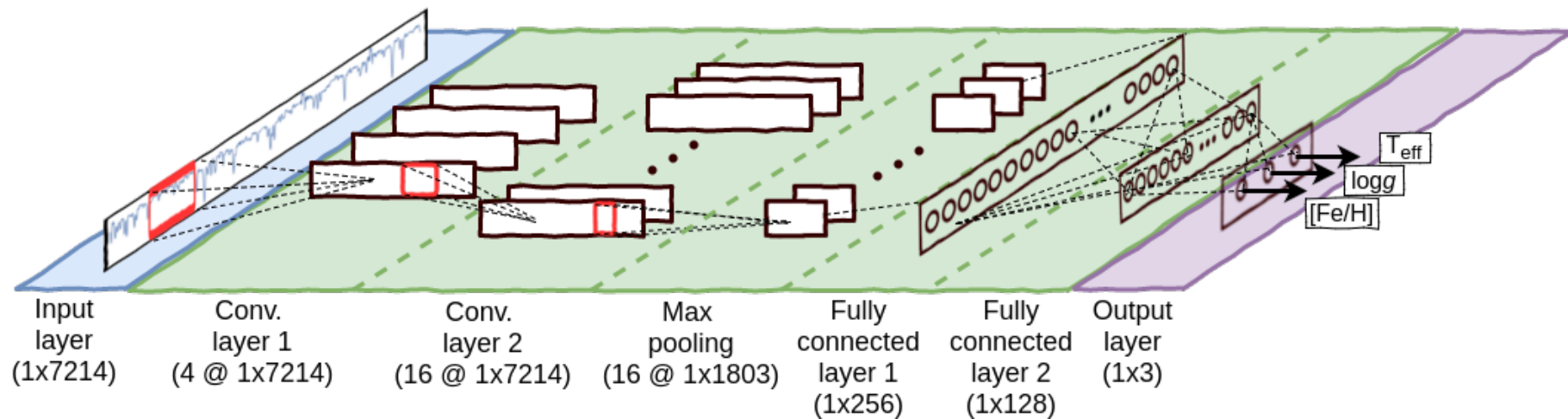


**distant past ML knowledge**

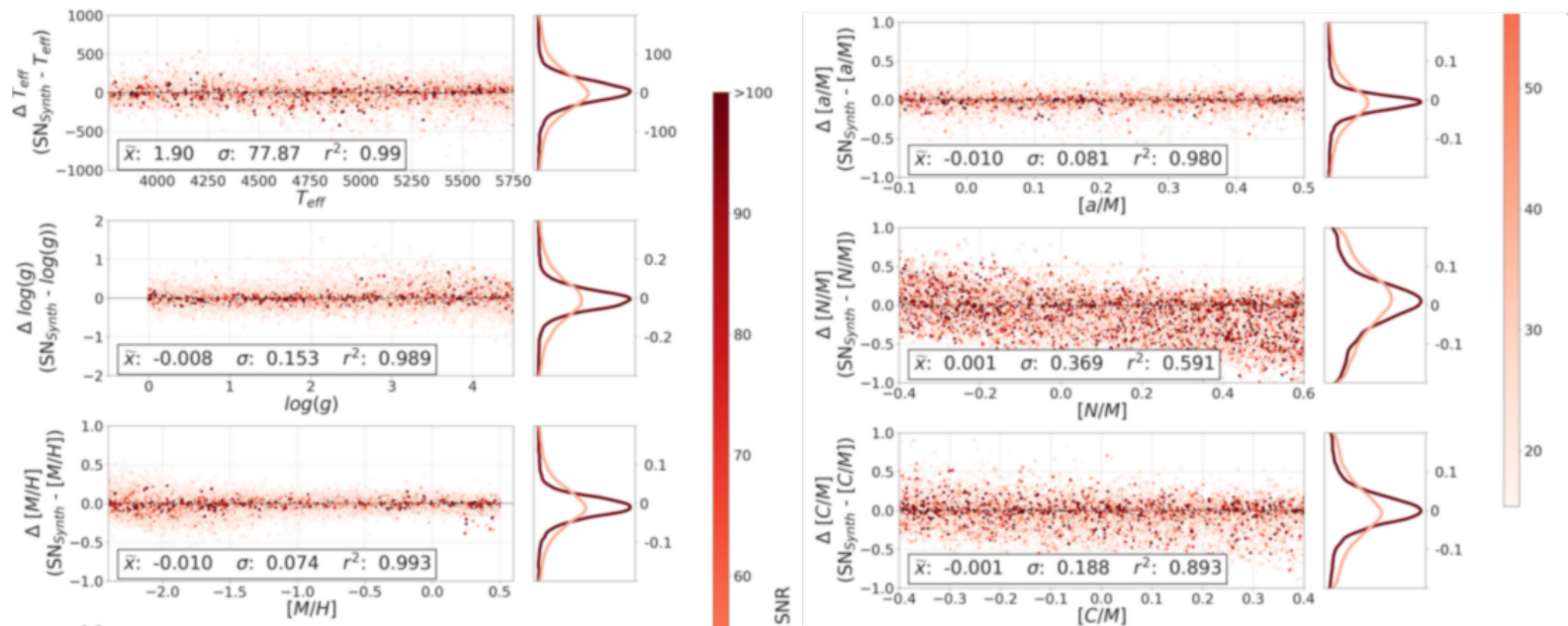
# stacking neural networks



# StarNet



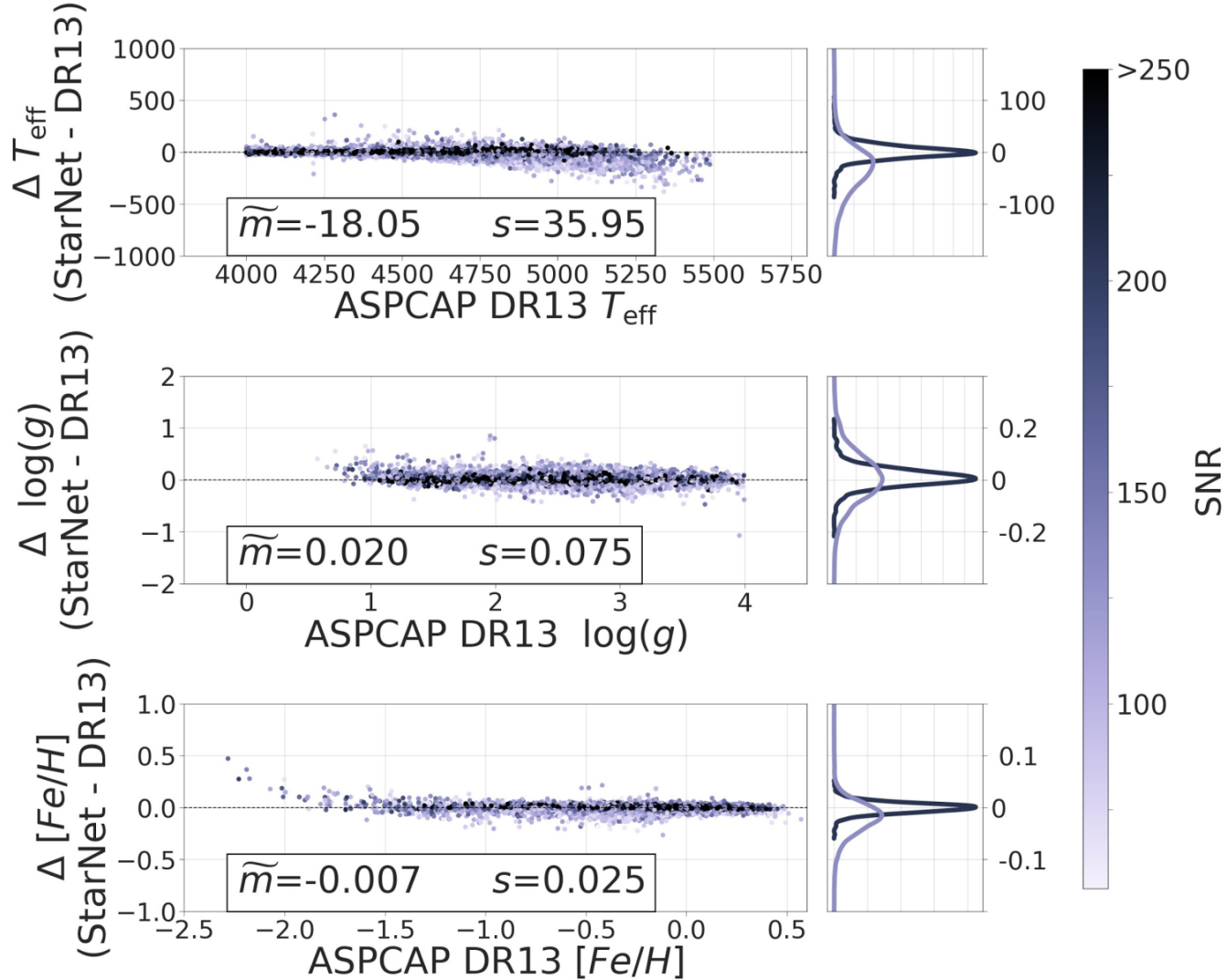
# starnet trained on synthetic spectra



# starnet trained on APOGEE spectra

15k stars with 47k visits

official pipeline parameters

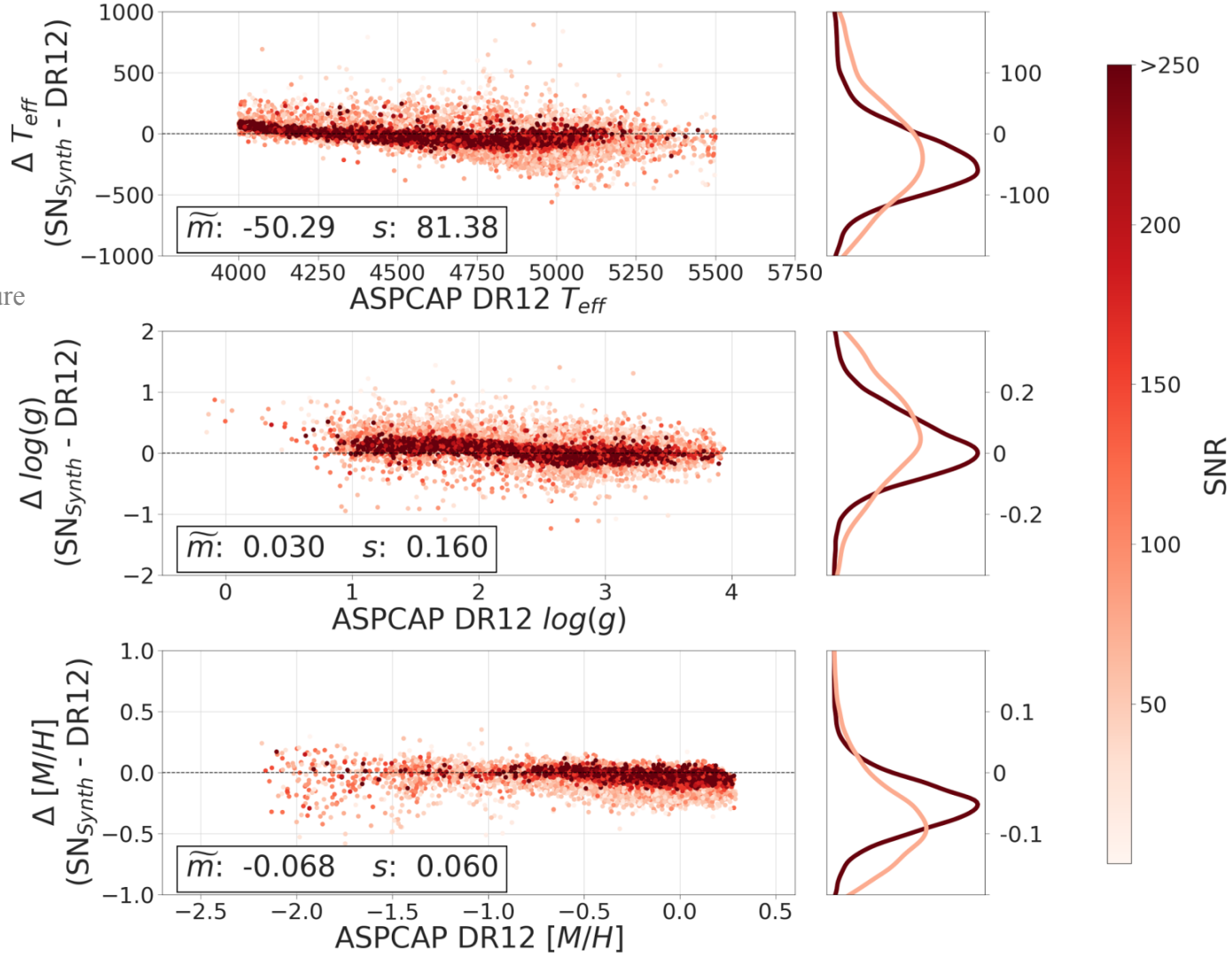




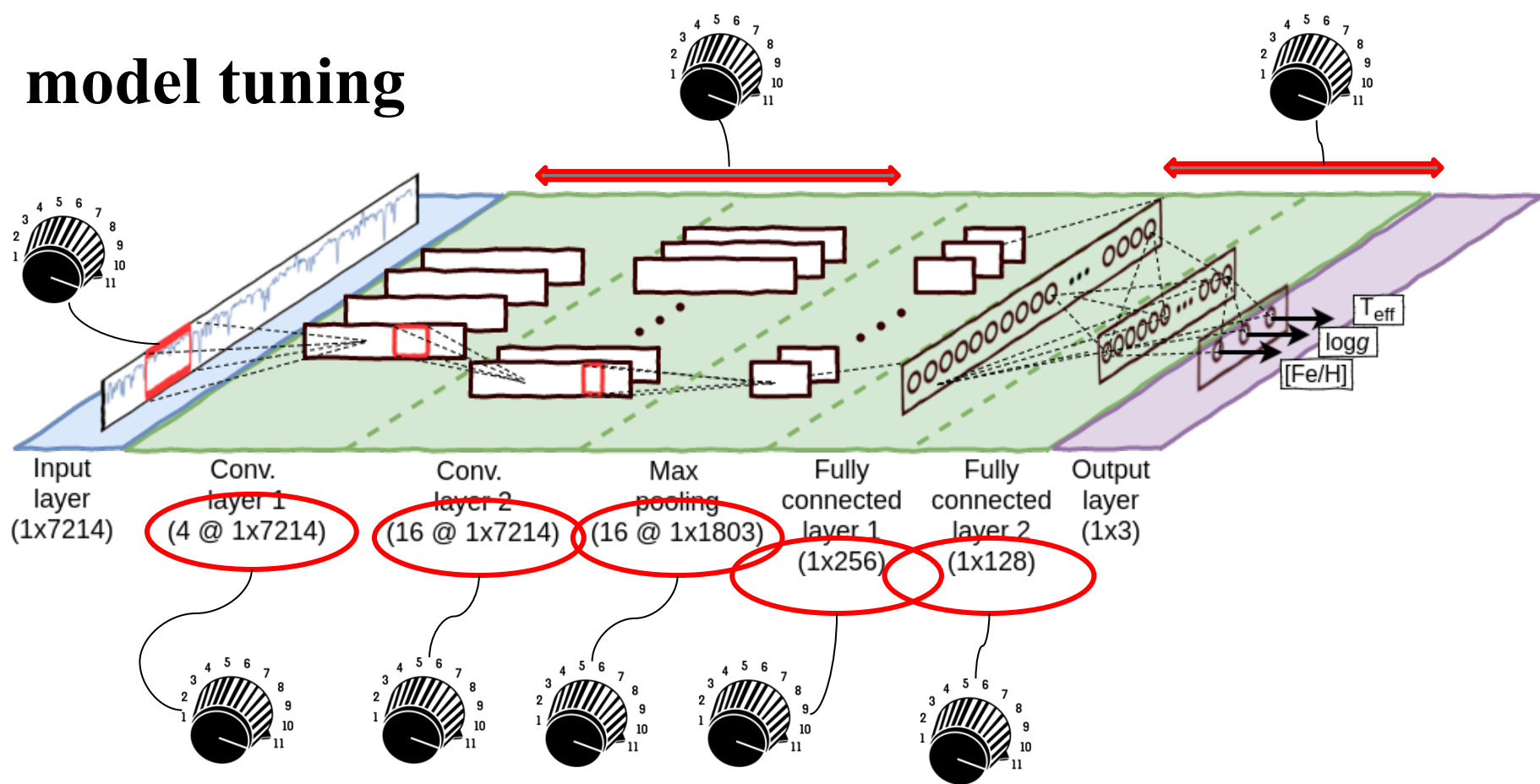
# trained on synth test on APOGEE

no modification to StarNet architecture

same 224k stars as simulations



# model tuning



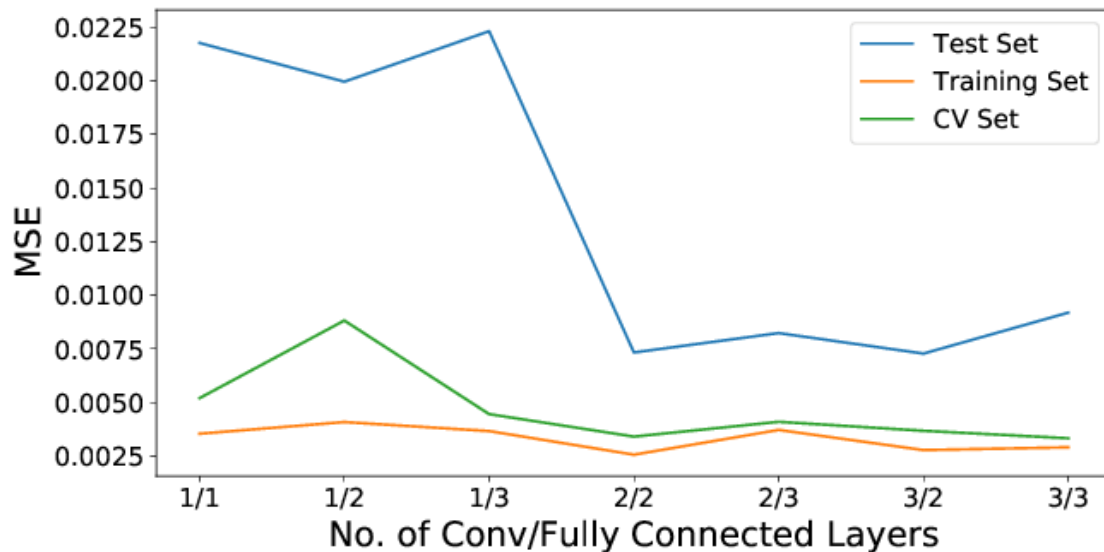
# meta learning starnet

13 hyperparameters

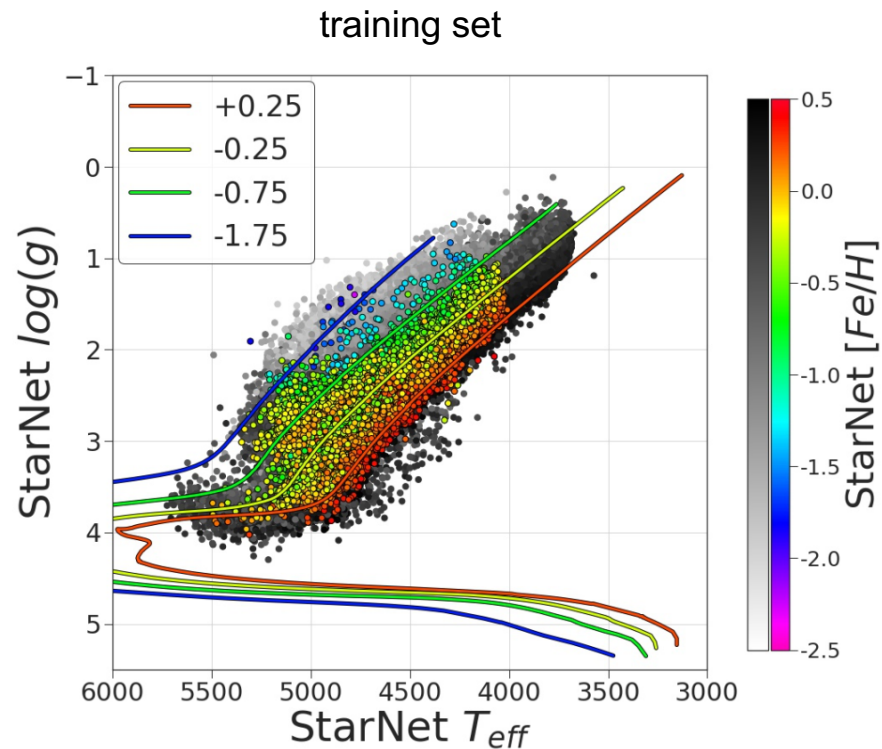
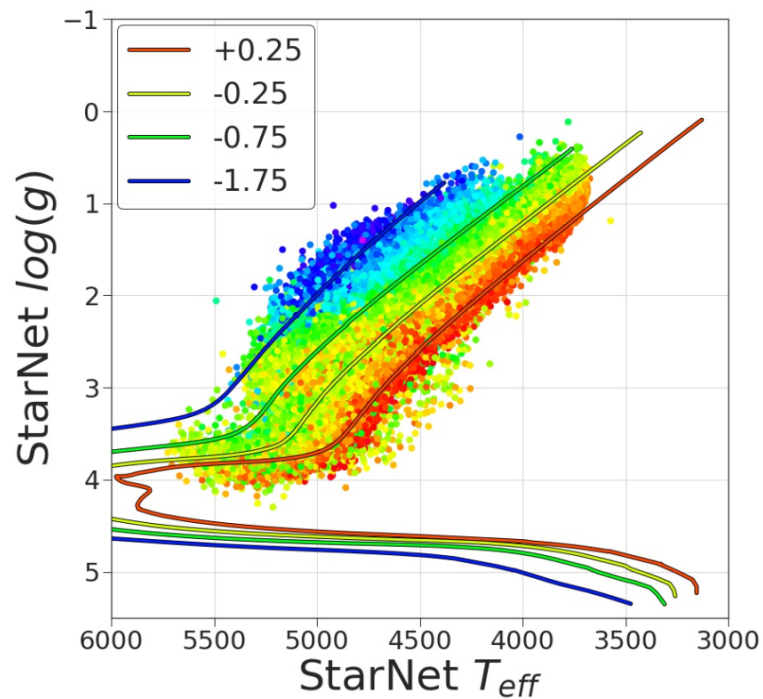
30mn per training (no GPU)

Tree-structured Parzen  
Estimator Bayes Optimizer

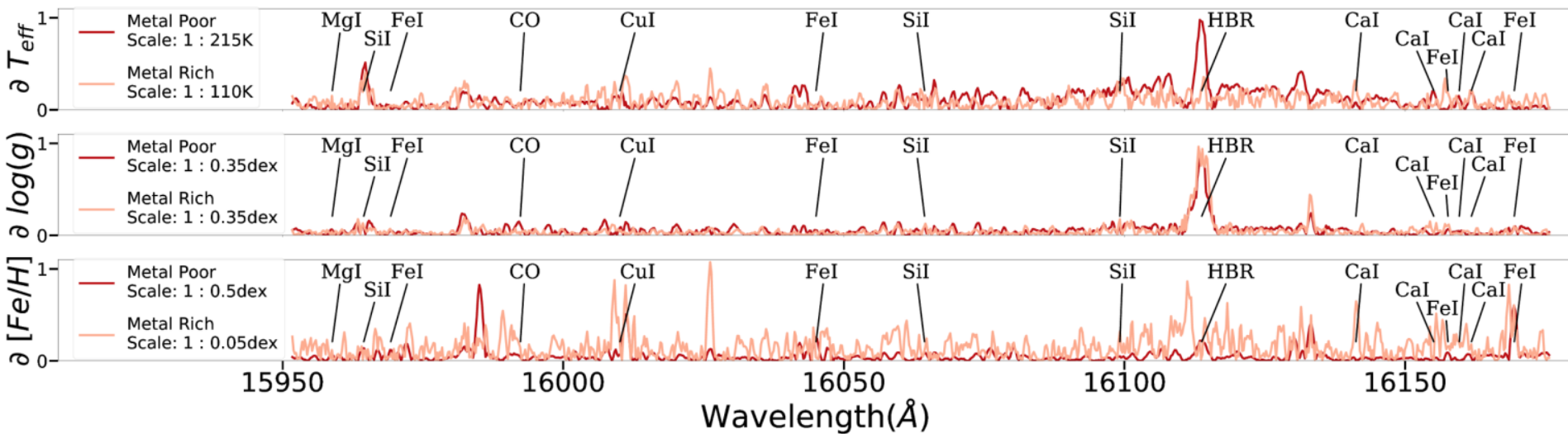
(smartish gridding)



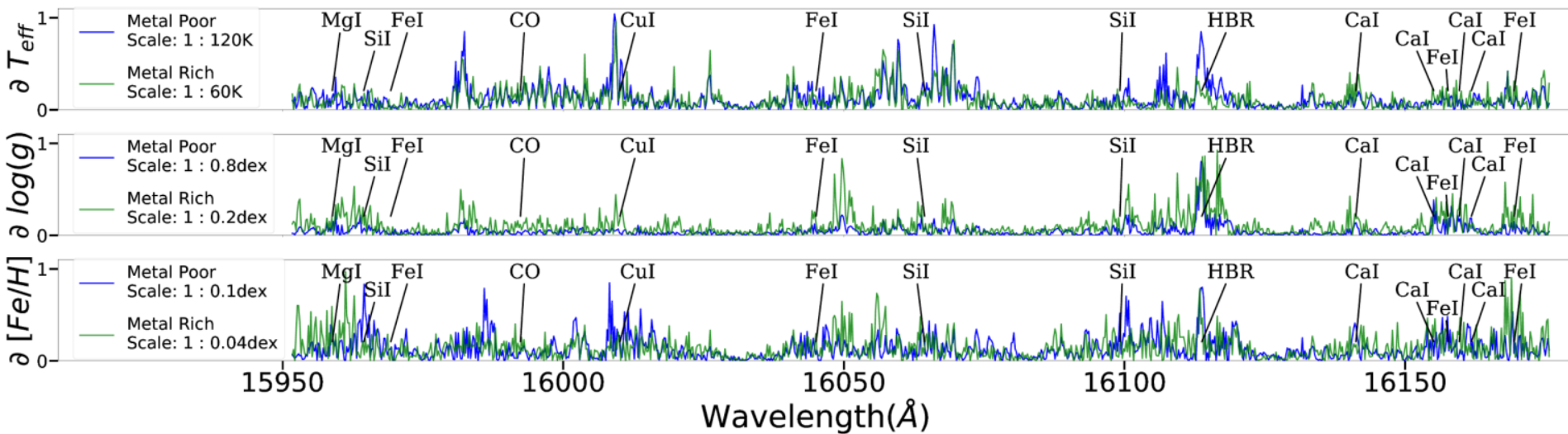
# realistic parameters



# where to look in a spectrum - synthetic

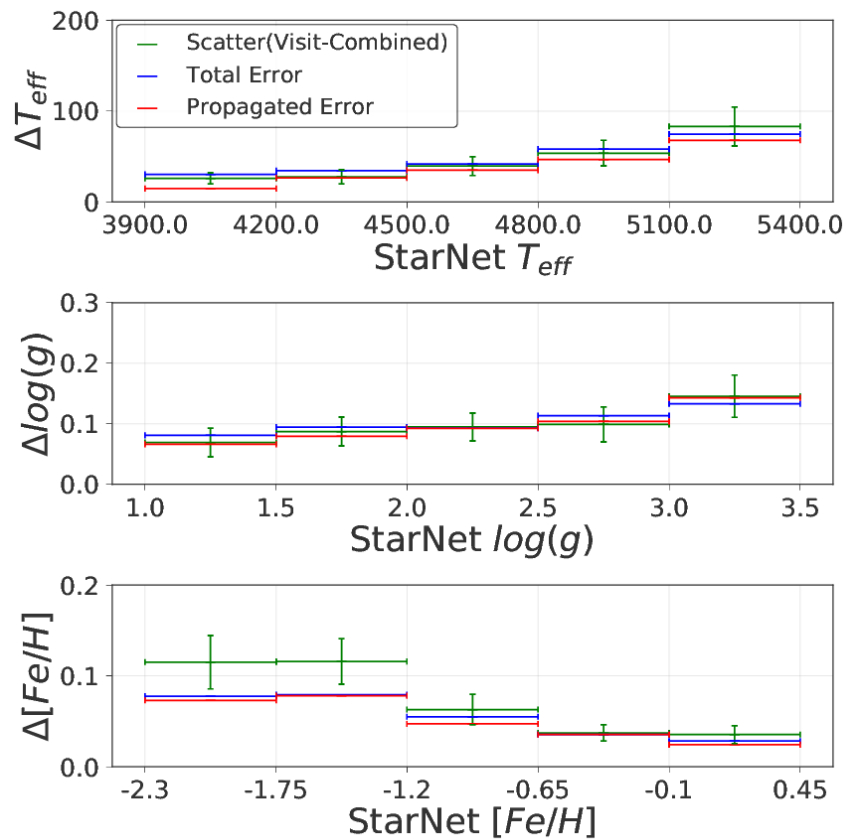


# where to look in a spectrum - real

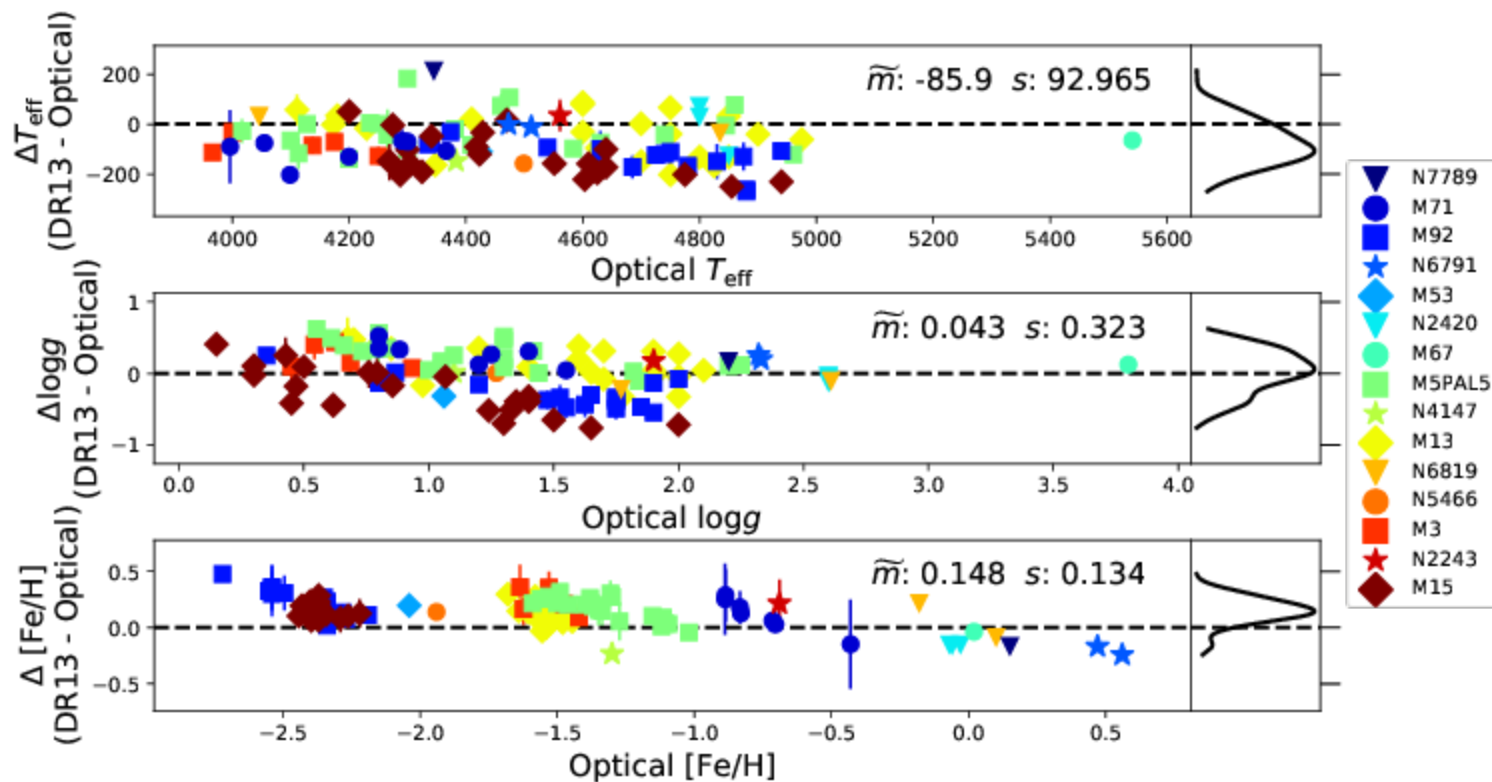


# uncertainties

propagated from error spectra onto the  
starnet model with intrinsic scatter.

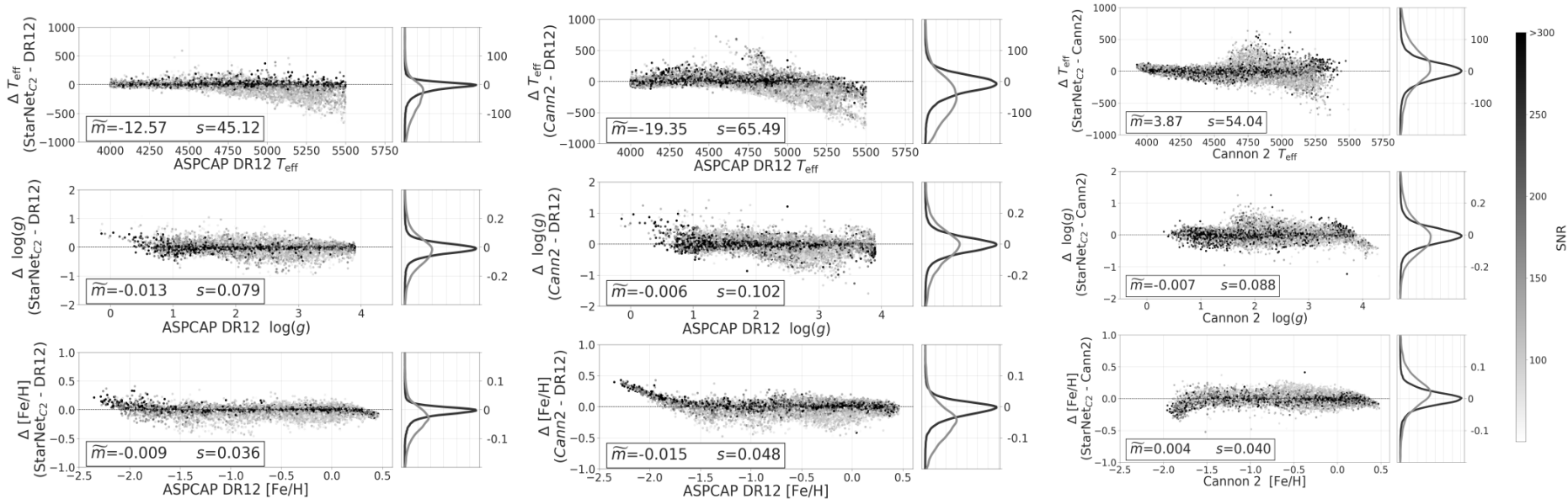


# calibration uncertainties





# comparison with official APOGEE methods






StarNet vs. ASCAP

The Cannon 2 vs. ASCAP

StarNet vs. The Cannon 2

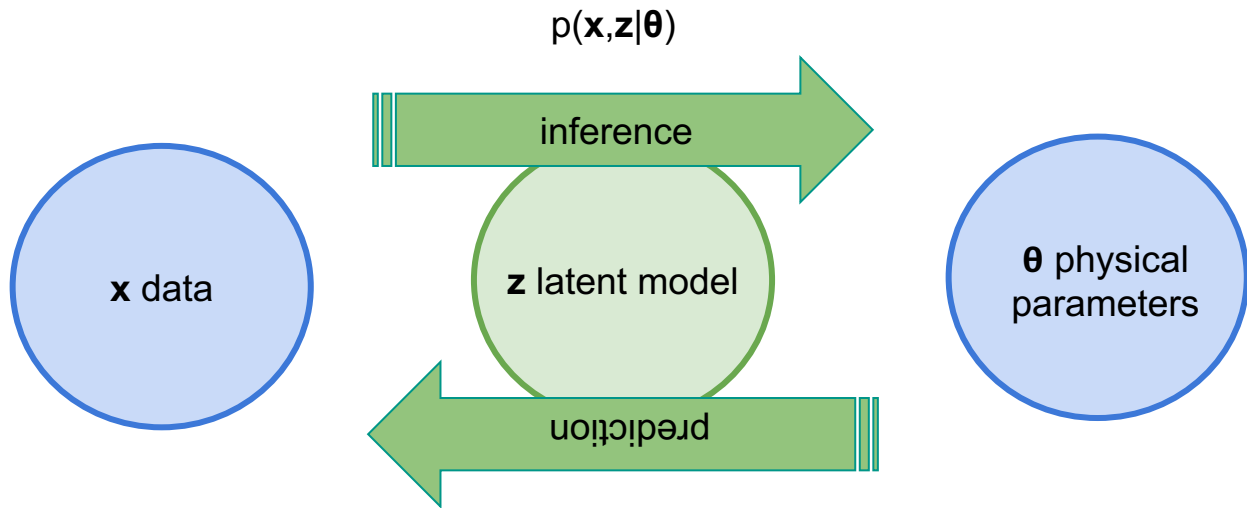
# machine learning limitations

- ~~dealing with uncertainties~~
- ~~dealing with heteroscedastic data~~
- ~~dealing with missing data~~
- interpretable models 
- generalisability 
- reduce training set size 

# machine learning limitations

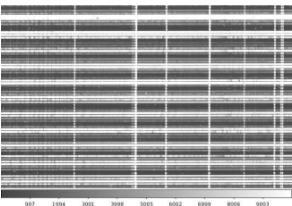
- dealing with uncertainties
- dealing with heteroscedastic data
- dealing with missing data
- interpretable models
- generalisability
- reduce training set size

# machine learning meets statistics

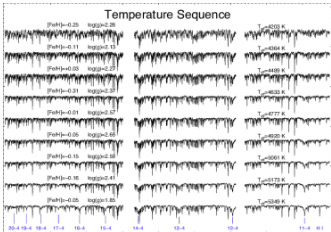


generative implicit models, deep probabilistic programming

raw 2D image



calibrated 1D spectra

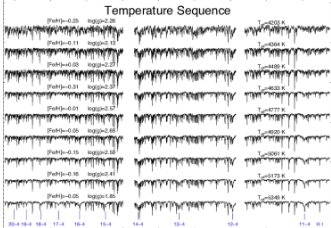


estimated stellar parameters

$T_{\text{eff}}$ ,  $\log(g)$ ,  $[\text{Fe}/\text{H}]$ ,  
 $[\text{C}/\text{H}]$ ,  $[\text{N}/\text{H}]$ ,  $[\text{O}/\text{H}]$ ,  
 $[\text{Na}/\text{H}]$ ,  $[\text{Mg}/\text{H}]$ ,  $[\text{Al}/\text{H}]$ ,  
 $[\text{Si}/\text{H}]$ ,  $[\text{S}/\text{H}]$ ,  $[\text{K}/\text{H}]$ ,  
 $[\text{Ca}/\text{H}]$ ,  $[\text{Ti}/\text{H}]$ ,  $[\text{V}/\text{H}]$ ,  
 $[\text{Mn}/\text{H}]$ ,  $[\text{Ni}/\text{H}]$ ,...

NTCO 2018?

synthetic 1D spectra



stellar parameters

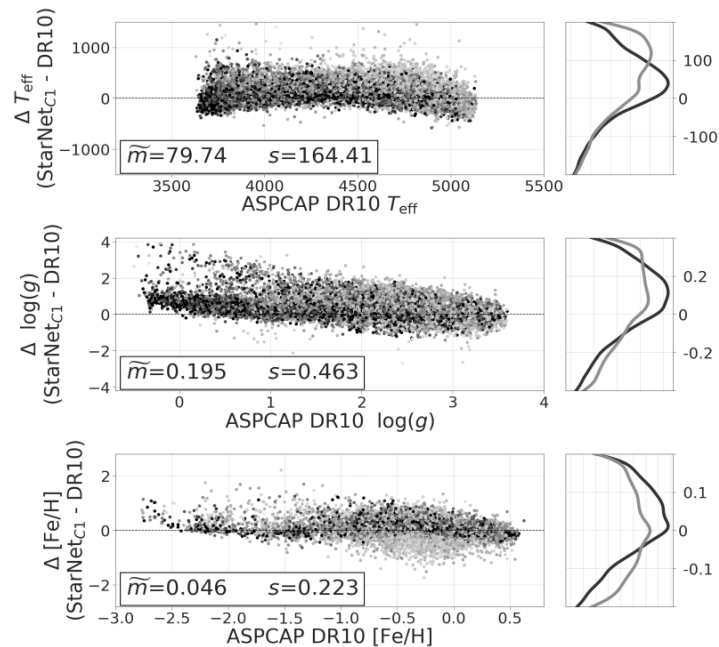
$T_{\text{eff}}$ ,  $\log(g)$ ,  $[\text{Fe}/\text{H}]$ ,  
 $[\text{C}/\text{H}]$ ,  $[\text{N}/\text{H}]$ ,  $[\text{O}/\text{H}]$ ,  
 $[\text{Na}/\text{H}]$ ,  $[\text{Mg}/\text{H}]$ ,  $[\text{Al}/\text{H}]$ ,  
 $[\text{Si}/\text{H}]$ ,  $[\text{S}/\text{H}]$ ,  $[\text{K}/\text{H}]$ ,  
 $[\text{Ca}/\text{H}]$ ,  $[\text{Ti}/\text{H}]$ ,  $[\text{V}/\text{H}]$ ,  
 $[\text{Mn}/\text{H}]$ ,  $[\text{Ni}/\text{H}]$ ,...

stellar populations  
galactic archeology  
near-field cosmology

**end**

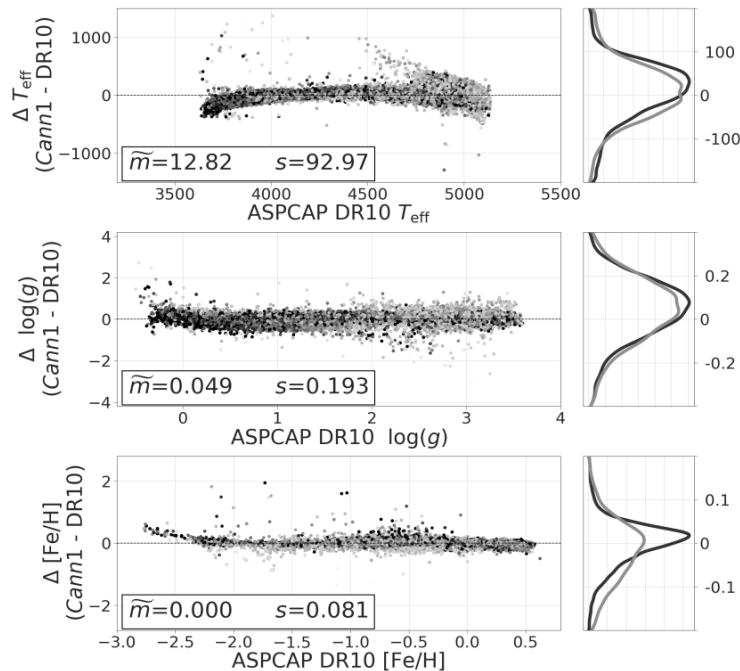
**extra slides**

# small training set



StarNet

542 DR10 stars

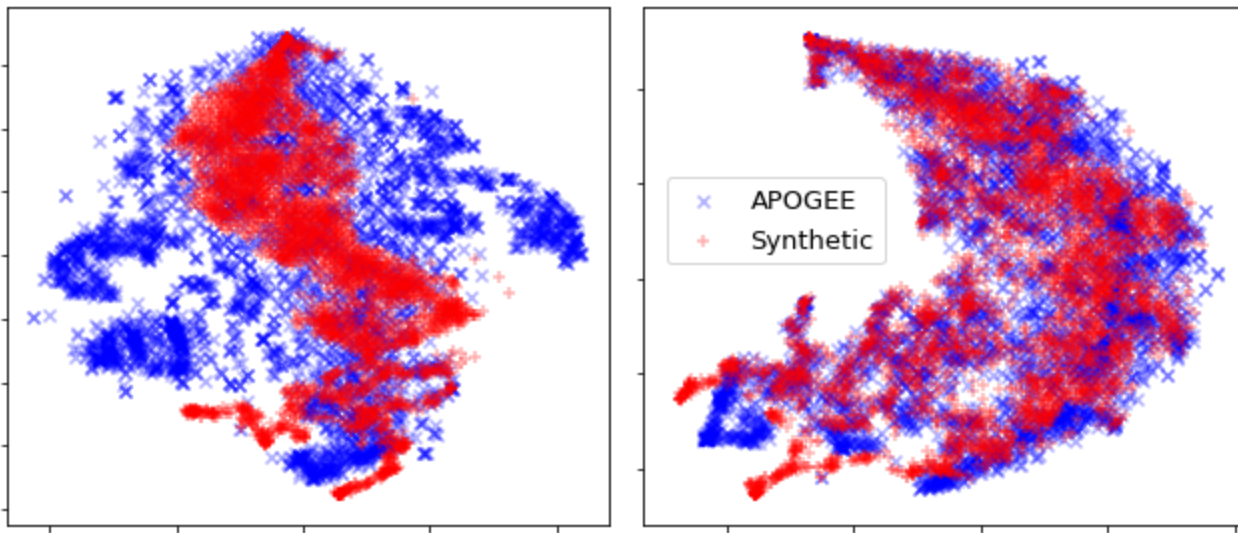


The Cannon Ness (2015)



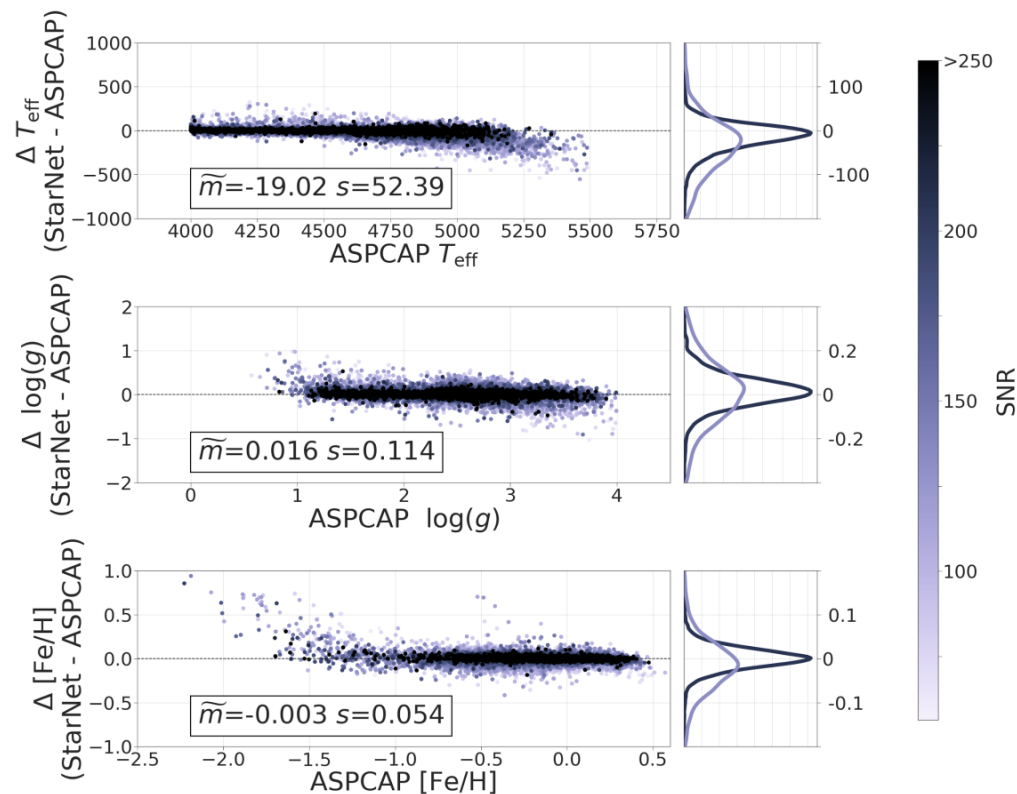
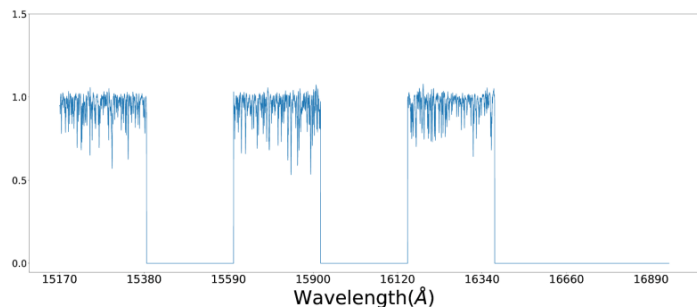
# synthetic gap?

t-SNE before and after zero-data - interpolation



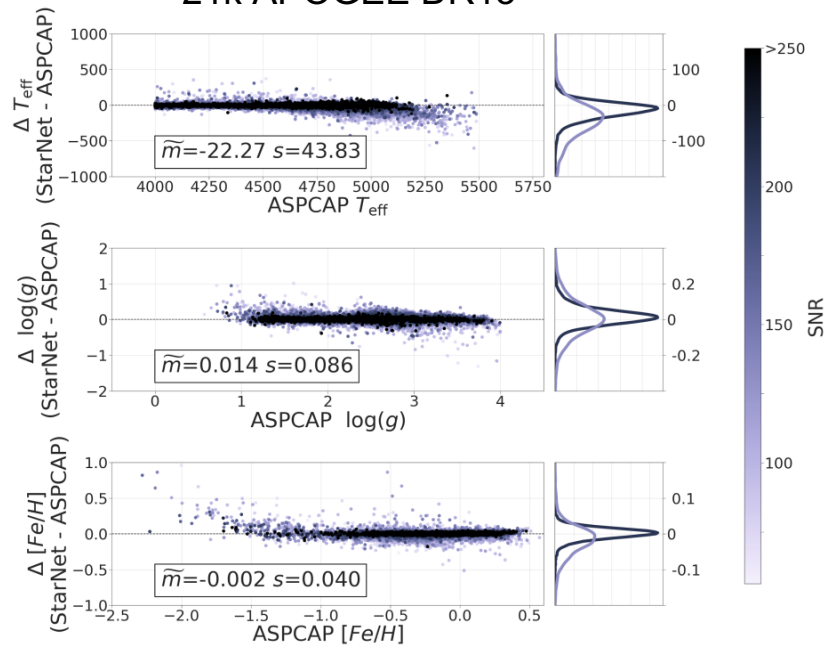
# missing data?

training and testing using random parts of the spectra

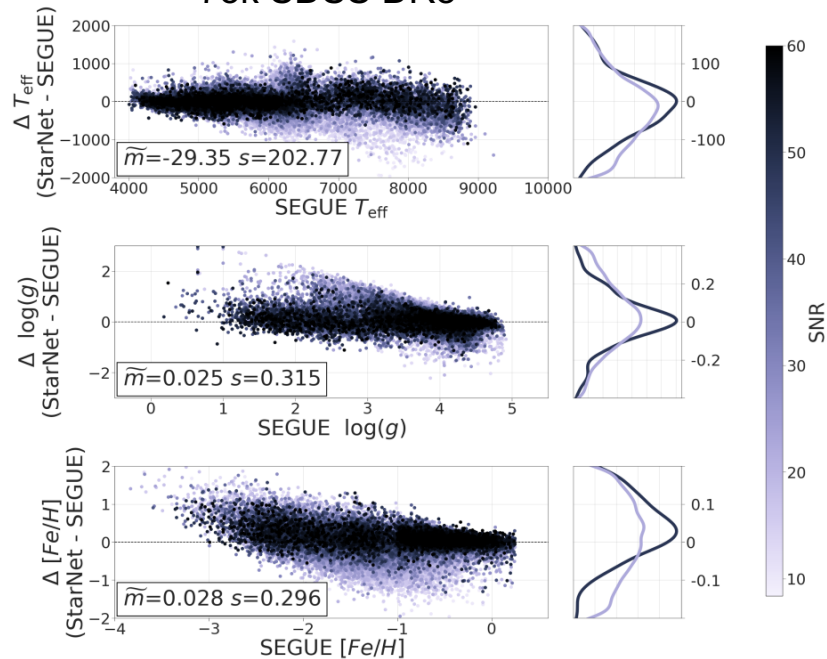


# multi-surveys

21k APOGEE DR13



75k SDSS DR8



simultaneous training on 41k IR with APSCAP + 77k optical spectra with SEGUE

# exploring deep learning architectures

Neural Network	no. of Filters in Conv. Layers	no. of Nodes in Fully Connected Layers	$r^2$	$T_{eff}$		$\log(g)$		$[Fe/H]$	
				$\tilde{x}$	$\sigma$	$\tilde{x}$	$\sigma$	$\tilde{x}$	$\sigma$
Shallow NN	0	512, 128	0.9667	-7.2	72.3	0.003	0.142	-0.008	0.057
Deep NN	0	2048, 1024, 512, 256, 128, 32	0.9642	-10.2	73.9	0.027	0.144	-0.011	0.057
Shallow CNN	16	128	0.9573	-1.2	82.9	0.017	0.155	-0.008	0.066
StarNet CNN	4, 16	256, 128	0.9749	-12.1	63.5	0.005	0.108	-0.014	0.049
Deep CNN	16, 32, 32, 64, 64	1024, 512, 256	0.9737	-8.7	70.2	0.003	0.105	-0.010	0.053